

"Fine Structure" of Hydrogen like atoms.

Let's first recall the spectrum of the

Hamiltonian $H_0 = \frac{p^2}{2m} - \frac{e^2}{r}$ which

corresponds to non-relativistic hydrogen atom.

The eigenstates are labelled as

$|n \ell m_\ell\rangle \otimes |m_s\rangle$ where 'n' is the
principle quantum number, $\ell = 0, 1, \dots, n-1$
is the orbital angular momentum,
' m_ℓ ' is the z-component of the
orbital angular momentum ($m_\ell = -\ell, -\ell+1, \dots, \ell$) and $m_s = \pm \frac{1}{2}$ is
the z-component of the intrinsic spin.

The eigenenergy of state $|n \ell m_\ell m_s\rangle$ depends only on 'n' and is given by $E_{n\ell m_\ell m_s}^0 = - \frac{mc^2 \alpha^2}{2n^2}$ where

$$\alpha = \left[\frac{e^2}{4\pi\epsilon_0 \hbar c} \right] \approx \frac{1}{137} \quad \text{is the fine-}$$

structure constant, and $mc^2 \approx 0.5 \text{ MeV}$ is the rest mass energy of the electron.

Therefore, the energy levels are hugely degenerate. For example, when $n=2$, there are six levels corresponding to $|n=2 \ell=1 m_\ell\rangle \otimes |m_s\rangle$ and two levels corresponding to $|n=2 \ell=0 m_\ell=0\rangle \otimes |m_s\rangle$, yielding a degeneracy of eight.

Since $\alpha \approx \frac{1}{137}$ is a small number,

one can imagine perturbations to H_0 that are proportional to various subleading powers of α , compared to H_0 . It turns out that at the leading order, relativistic corrections are of order α^4 . A satisfactory treatment requires studying quantum electrodynamics (QED) / Dirac equation. Here we will only pursue a heuristic treatment.

There are two separate corrections of $O(\alpha^4)$, one due to relativistic dispersion, and other due to spin-

orbit coupling.

(a) α^4 correction due to relativistic dispersion:

The relativistic Hamiltonian, ignoring spin, is $H = \sqrt{p^2 c^4 + m^2 c^4} - mc^2 - \frac{e^2}{r}$

$$\simeq \underbrace{\frac{p^2}{2m} - \frac{e^2}{r}}_{H_0} - \underbrace{\frac{p^4}{8m^3 c^2}}_{\Delta H_1}.$$

Therefore, the correction to the energy of a level $|n \ell m_\ell\rangle \otimes |m_s\rangle$ is given by $\delta E_{n\ell m_\ell m_s}^{(1)} = - \langle n \ell m_\ell m_s | \frac{p^4}{8m^3 c^2} | n \ell m_\ell m_s \rangle$

To evaluate this expectation value, one may rewrite it as

$$\delta E_{nlm}^{(1)} = -\frac{1}{2mc^2} \left\langle \left(\frac{p^2}{2m} \right)^2 \right\rangle$$

where $\langle \rangle$ denotes expectation value w.r.t. $|nlm\rangle$

$$= -\frac{1}{2mc^2} \left\langle \left(\frac{e^2}{r} + H_0 \right)^2 \right\rangle$$

$$= -\frac{1}{2mc^2} \left\langle H_0^2 + H_0 \frac{e^2}{r} + \frac{e^2}{r} H_0 + \frac{e^4}{r^2} \right\rangle$$

$$= -\frac{1}{2mc^2} \left[(E_n^0)^2 + 2E_n^0 \left\langle \frac{e^2}{r} \right\rangle + e^4 \left\langle \frac{1}{r^2} \right\rangle \right]$$

One may use virial theorem to show $\left\langle \frac{e^2}{r} \right\rangle = -2E_n^0$.

finally $\left\langle \frac{e^4}{r^2} \right\rangle = \frac{4n(E_n^0)^2}{2 + \frac{1}{2}}$

[see Gottfried/Shankar for details].

Combining everything,

$$\delta E_{\text{nlmms}} = -\frac{1}{2mc^2} \left[-3(E_n^0)^2 + \frac{4n(E_n^0)^2}{2 + \frac{1}{2}} \right]$$

Using $E_n^0 = -\frac{\alpha^2 mc^2}{2n^2}$

$$\Rightarrow \delta E_{\text{nlmms}}^{(1)} = -\frac{mc^2}{2} \alpha^4 \left[-\frac{3}{4n^4} + \frac{1}{n^3(2 + \frac{1}{2})} \right]$$

(b) α^4 correction due to spin-orbit coupling:

Classically, the electron experiences a magnetic field due to relative motion w.r.t. the proton:

$$\vec{B} = -\frac{e}{c} \frac{\vec{v} \times \vec{r}}{r^3}$$

Since electron carries an intrinsic magnetic moment $\vec{\mu} (\propto \vec{S})$ where $\vec{S}$ is the intrinsic spin, the coupling between $\vec{B}$ and $\vec{\mu}$ leads to a correction to H_0 of the form:

$$\begin{aligned}\Delta H_2 &= -\vec{\mu} \cdot \vec{B} \\ &= \frac{e \vec{\mu} \cdot (\vec{p} \times \vec{r})}{mc r^3}\end{aligned}$$

$$= \frac{-e \vec{\mu} \cdot \vec{L}}{mc r^3}$$

where $\vec{L} = \vec{r} \times \vec{p}$ is the orbital angular momentum. Classically, the

proportionality constant between $\vec{\mu}$

and $\vec{S}$ is given as $\vec{\mu} = \frac{-e}{mc} \vec{S}$

(see, e.g. Griffiths Electrodynamics

Chapter 5). However, quantum mechanically

there is an additional factor of $\frac{1}{2}$

which can be derived using Dirac

equation. Putting everything together,

$$\Delta H_2 = \frac{e^2}{2m^2 c^2 r^3} \vec{S} \cdot \vec{L}$$

Within degenerate level perturbation

theory, $H_{\text{eff}} = H_0 + \frac{e^2}{2m^2c^2r^3} \vec{S} \cdot \vec{L}$
+ higher order terms

and therefore, one needs to diagonalize

the perturbation $\vec{S} \cdot \vec{L}$. Since

$$\vec{S} \cdot \vec{L} = \frac{1}{2} [(\vec{S} + \vec{L})^2 - S^2 - L^2],$$

the eigenstates of $\vec{S} \cdot \vec{L}$ are

labelled by $\vec{J}^2 = (\vec{S} + \vec{L})^2$, L^2
 $= l(l+1)$ and $J_z = L_z + S_z$.

Using angular momentum addition

rules, there are only two

possibilities: $\vec{J}^2 = \hat{j}(\hat{j}+1)$ with

$$\hat{j} = l + \frac{1}{2} \quad \text{or} \quad \hat{j} = l - \frac{1}{2}$$

when $j = l + \frac{1}{2}$,

$$H_{\text{eff}} = H_0 + \frac{e^2}{2m^2 c^3 r^3} \left[\left(l + \frac{1}{2}\right) \left(l + \frac{3}{2}\right) - l(l+1) - \frac{3}{4} \right]$$

$$= H_0 + \frac{e^2}{2m^2 c^3} \frac{l}{r^3}$$

when $j = l - \frac{1}{2}$,

$$H_{\text{eff}} = H_0 + \frac{e^2}{2m^2 c^3 r^3} \left[\left(l - \frac{1}{2}\right) \left(l + \frac{1}{2}\right) - l(l+1) - \frac{3}{4} \right]$$

$$= H_0 + \frac{e^2}{2m^2 c^3 r^3} [-l - 1]$$

The correction to energy will be

$$\delta E^{(2)} = \langle n l m_l m_s | \frac{1}{r^3} | n l m_l m_s \rangle \times \frac{e^2}{2m^2 c^3 r^3} \times \begin{cases} -l & \text{when } j = l + \frac{1}{2} \\ -l-1 & \text{when } j = l - \frac{1}{2} \end{cases}$$

One can show,

$$\left\langle \frac{1}{r^3} \right\rangle = \frac{1}{a_0^3 n^3 l(l+1)(l+\frac{1}{2})}$$

where a_0 is the Bohr radius.

⇒ The total correction to the energy at $O(\alpha^4)$ is given by

$$\delta E = \delta E^{(1)} \text{ [due to relativistic kinetic energy]}$$

$$+ \delta E^{(2)} \text{ [due to spin-orbit interaction]}$$

when $j = l + \frac{1}{2} :$

$$\delta E = -\frac{mc^2}{2} \alpha^4 \left[-\frac{3}{4n^4} + \frac{1}{n^3(l+\frac{1}{2})} \right]$$

$$+ \frac{mc^2}{4} \frac{\alpha^4}{n^3(l+\frac{1}{2})(l+1)}$$

$$= -\frac{mc^2}{2n^3} \alpha^4 \left[-\frac{3}{4n} + \frac{1}{l+\frac{1}{2}} - \frac{1}{2(l+\frac{1}{2})(l+\frac{1}{2})} \right]$$

$$= -\frac{mc^2}{2n^3} \alpha^4 \left[-\frac{3}{4n} + \frac{1}{l+1} \right]$$

when $j = l - \frac{1}{2} :$

$$\delta E = -\frac{mc^2}{2} \alpha^4 \left[-\frac{3}{4n^4} + \frac{1}{n^3(l+\frac{1}{2})} \right]$$

$$- \frac{mc^2}{4} \frac{\alpha^4}{n^3} \frac{1}{l(l+\frac{1}{2})}$$

$$= \frac{-mc^2}{2n^3} \alpha^4 \left[-\frac{3}{4n} + \frac{1}{l + \frac{1}{2}} + \frac{1}{2l(l + \frac{1}{2})} \right]$$

$$= \frac{-mc^2}{2n^3} \alpha^4 \left[-\frac{3}{4n} + \frac{1}{l} \right]$$

Therefore, in both cases ($j = l \pm \frac{1}{2}$),

$$\delta E = \frac{-mc^2}{2n^3} \alpha^4 \left[-\frac{3}{4n} + \frac{1}{j + \frac{1}{2}} \right]$$

Although we assumed that $l \neq 0$ (otherwise $\langle L \cdot S \rangle = 0$), the above expression is true even when $l = 0$ (so that $j = s = \frac{1}{2}$).